Sphere Rigidity in the Euclidean Space

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The well-known Alexandrov theorem [1] says that embedded hypersurfaces in \mathbb{R}^{n+1} with constant mean curvature are geodesic spheres. This result is not true for only immersed hypersurfaces. For instance, the so-called Wente's tori (see [5]) are examples of compact surfaces with constant mean curvature in \mathbb{R}^3 , which are not geodesic spheres. Other examples of higher genus are known (see [3] for instance).

For immersed hypersurfaces of constant mean curvature, an additional assumption is needed. One condition is given by the Hopf theorem [2], which says that constant mean curvature spheres immersed in \mathbb{R}^{n+1} are geodesic spheres.

In this talk, we give a new rigidity theorem for spheres, where we replace the topological assumption by a metric assumption. Precisely, it is easy to see that hypersurfaces of \mathbb{R}^{n+1} with constant mean curvature and constant scalar curvature are geodesic spheres. This result comes from the fact that a hypersurface of constant mean curvature and constant scalar curvature is totally umbilic. Here, we give a new rigidity result with a weaker assumption on the scalar curvature. Namely, we show

Théorème 1 Let (M^n, g) be a compact, connected and oriented Riemannian manifold without boundary isometrically immersed into \mathbb{R}^{n+1} . Let h be a positive constant. Then, there exists $\varepsilon > 0$ such that if

- (1) H = h and
- (2) $|\text{Scal} s| \leq \varepsilon$,

for some constant s, then M is the sphere $\mathbb{S}^n\left(\frac{1}{h}\right)$ with its standard metric.

We derive this theorem from results about almost umbilic hypersurfaces that we will prove before, and based on a previous eignevalue pinching result given in [4]. We will give a word about this eigenvalue pinching result.

References

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